

Laboratory Simulations of Aqueous Reactions in Lunar Polar Regions

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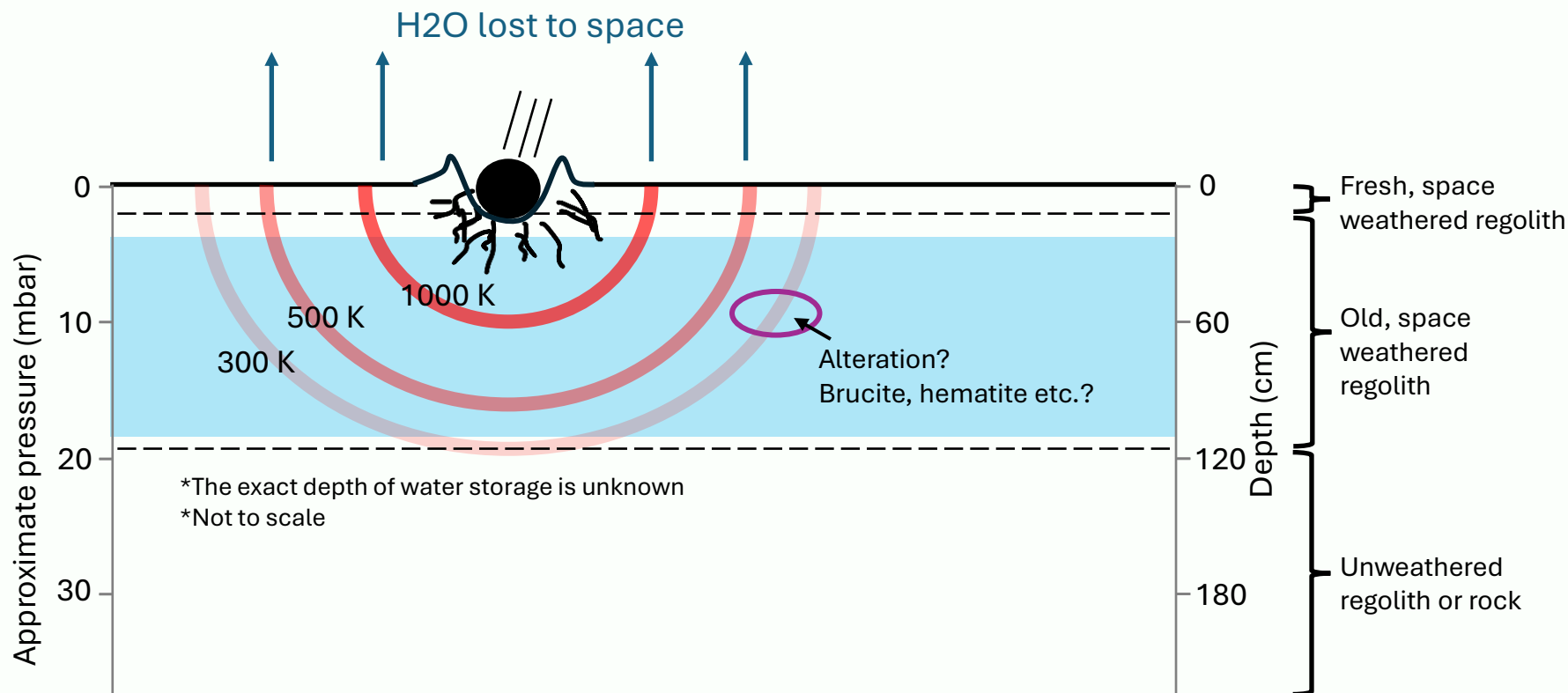


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& Planetary Sciences

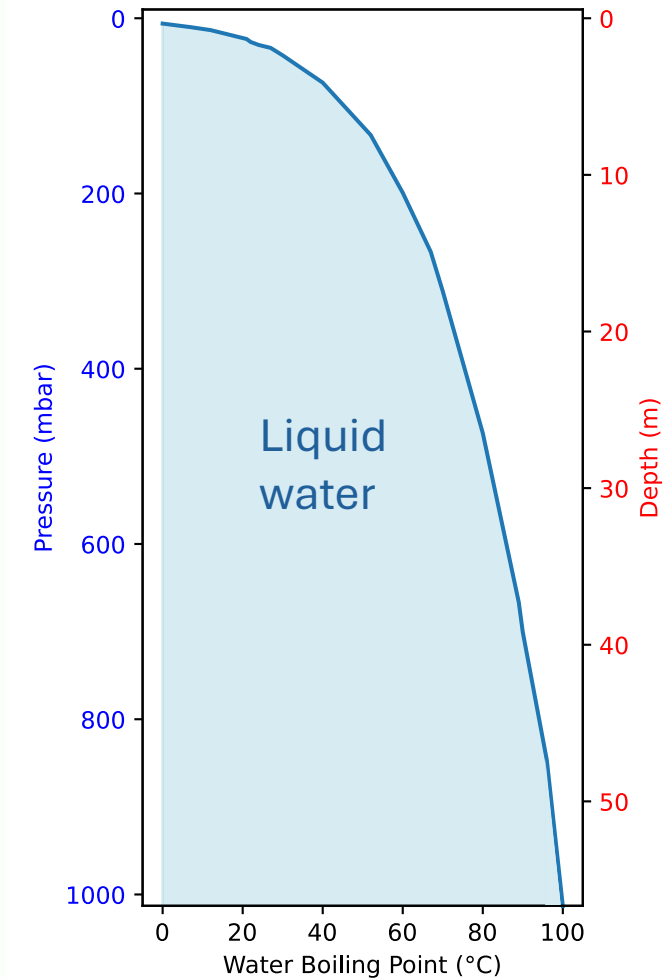
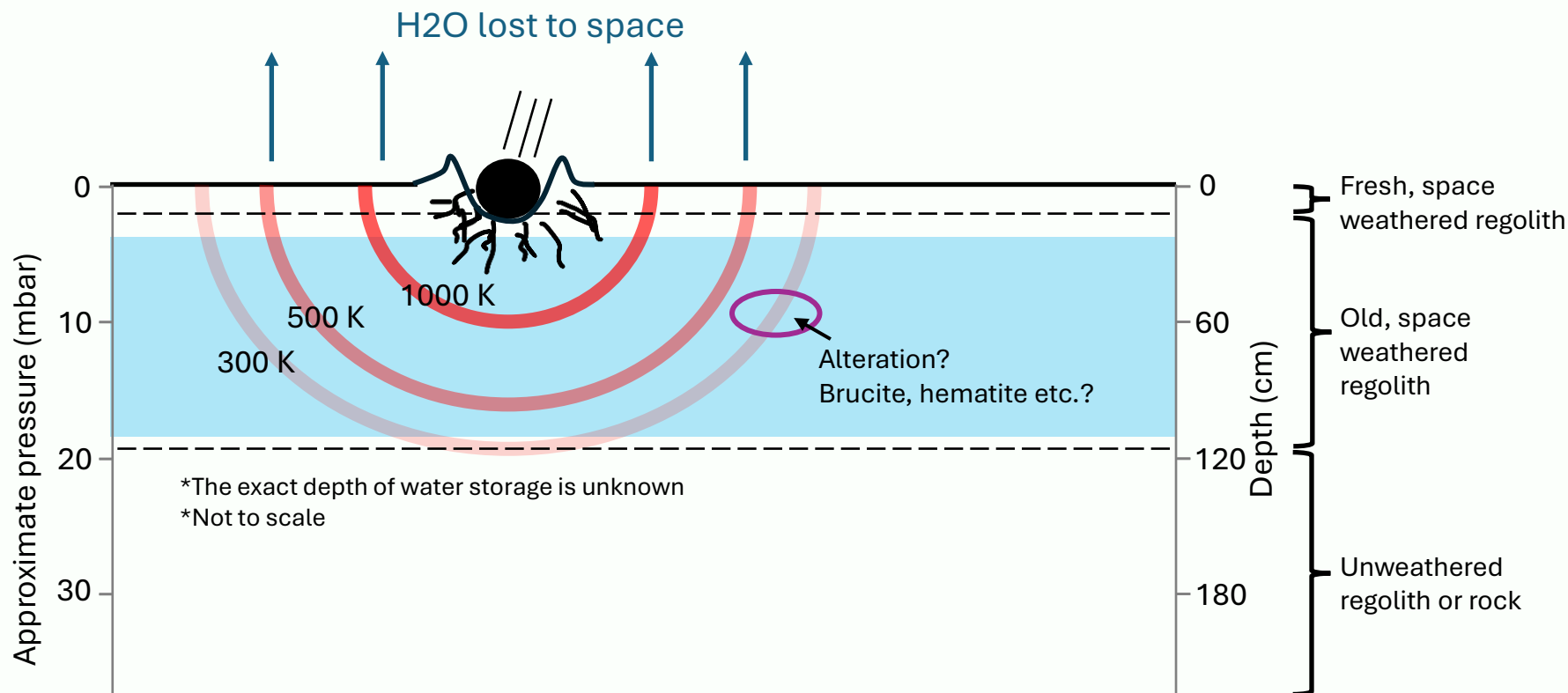
Background

- We might expect aqueous alteration on the Moon in the polar regions
- Heat + ice + pressure + long duration = alteration
- Experimental methods to simulate alterations



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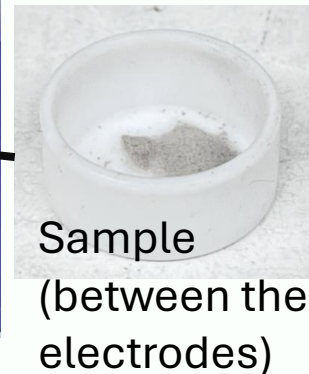
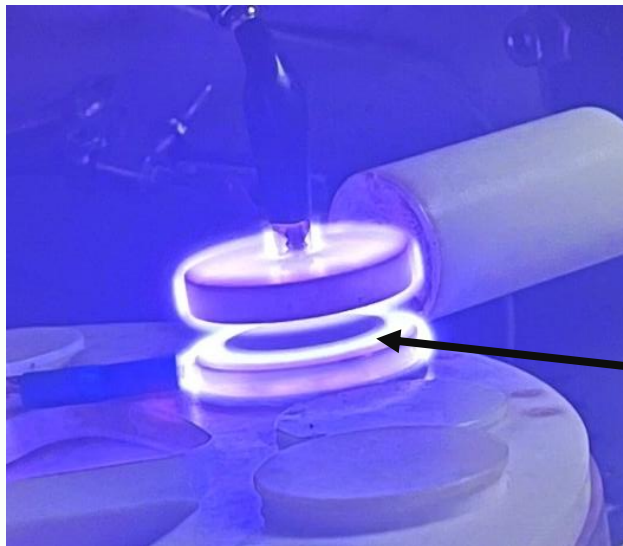


Experiment 1: Vacuum chamber + Energetic electron impact (EEI)

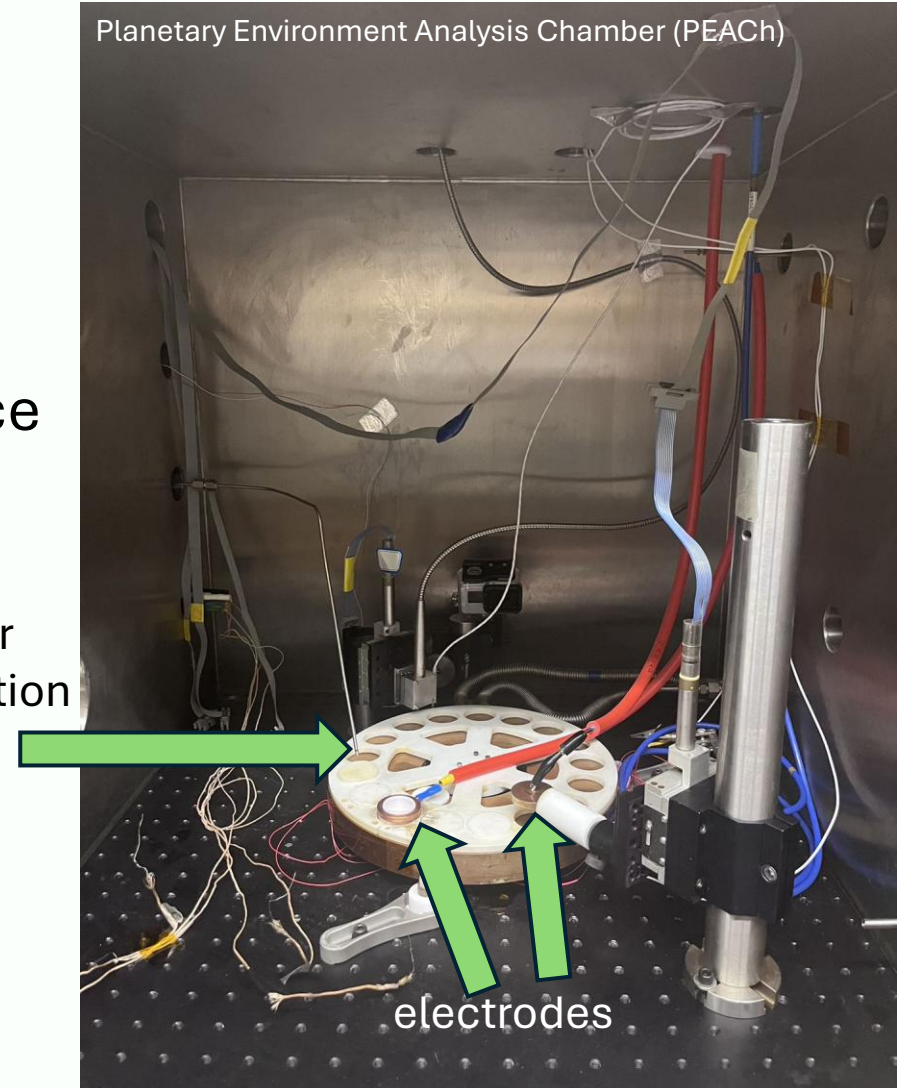
Key functions:

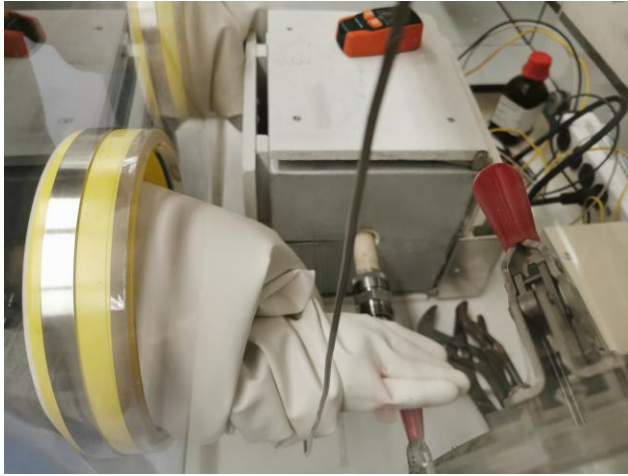
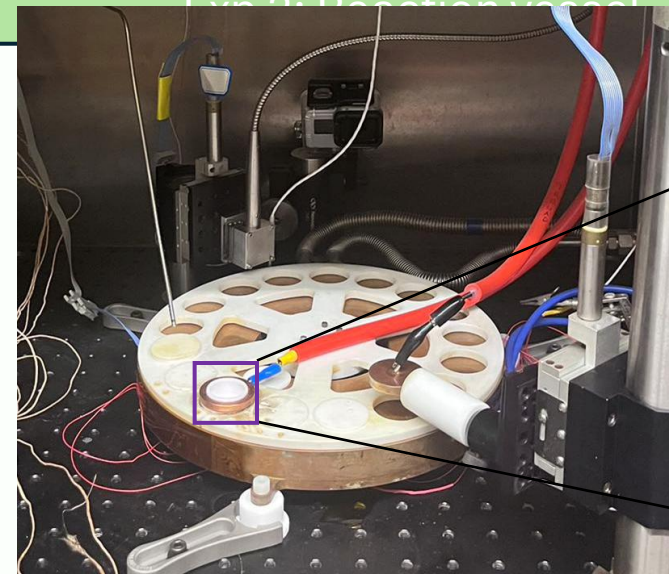
- Using EEI to activate the sample surface
- React water with olivine and basaltic glass
- Use D_2O (to distinguish from atmospheric H_2O)
- Controlled P-T condition ~16m under the surface

Energetic
electron
impact (EEI)
in Argon

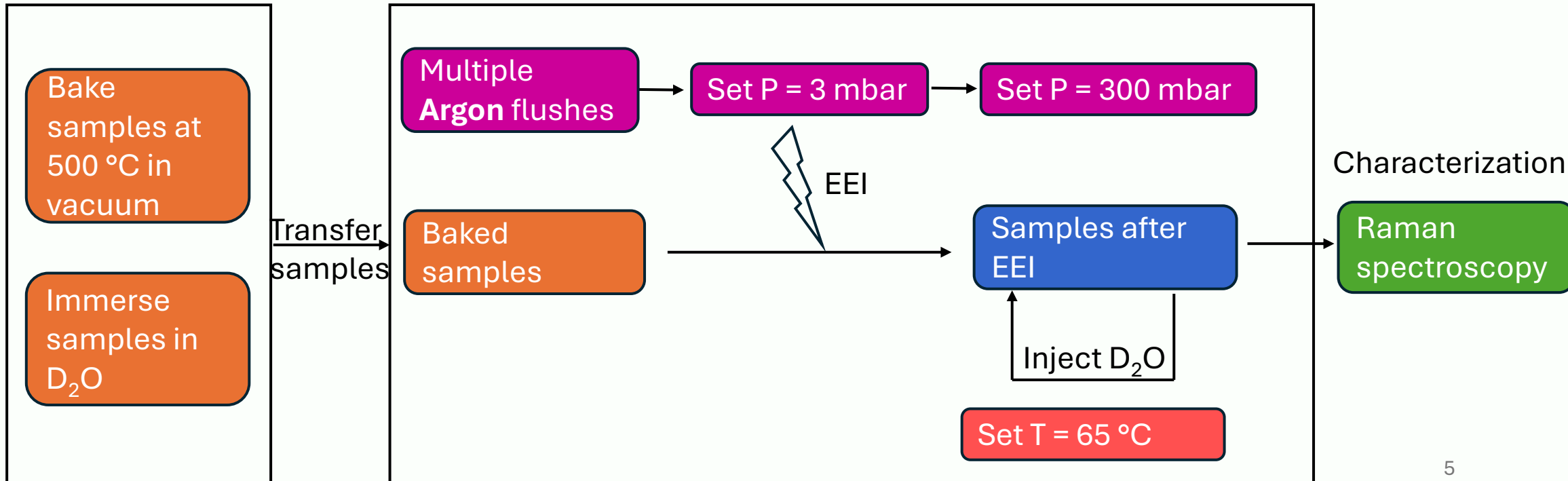


Water
injection
port

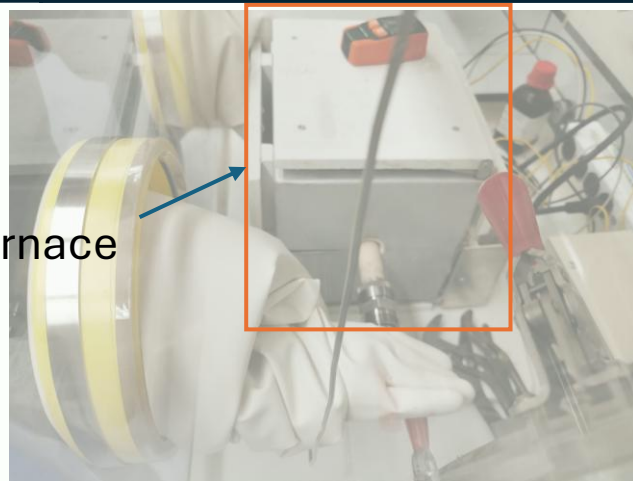


Glove box (N_2 gas)

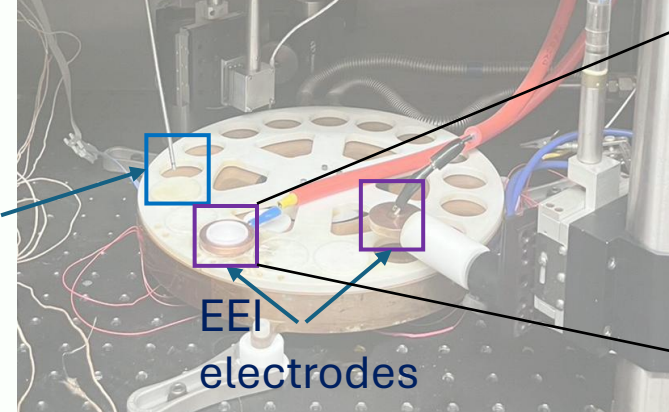
Vacuum chamber (Ar gas)

Sample
(olivine)

Furnace

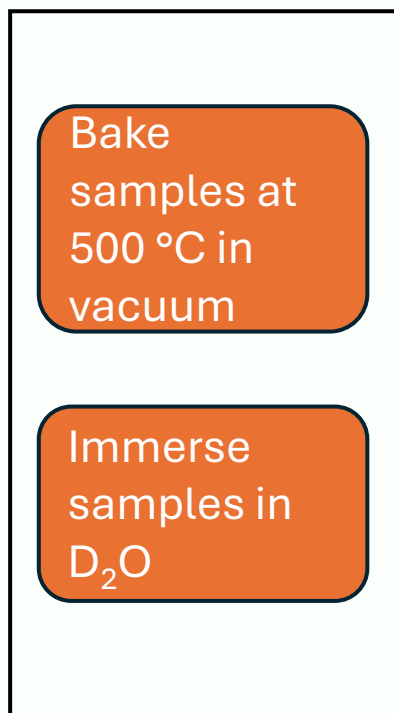


Injection port

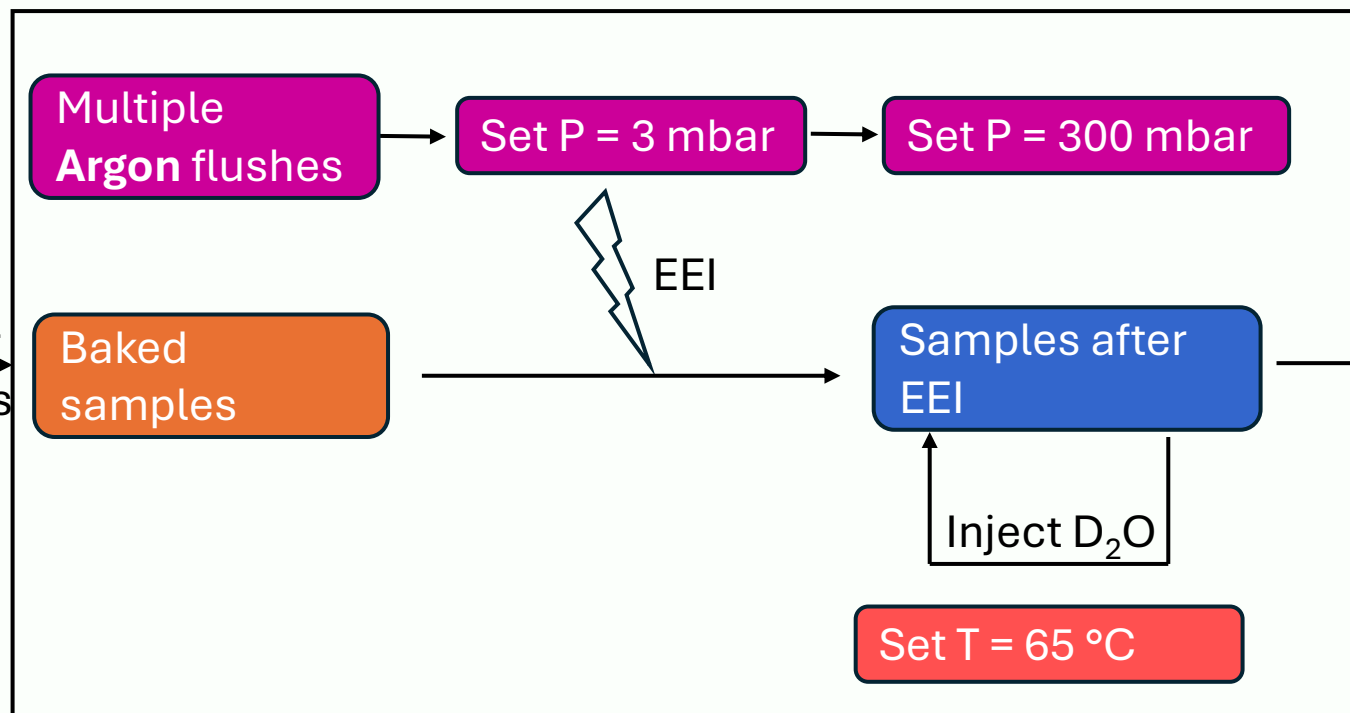


EEI electrodes

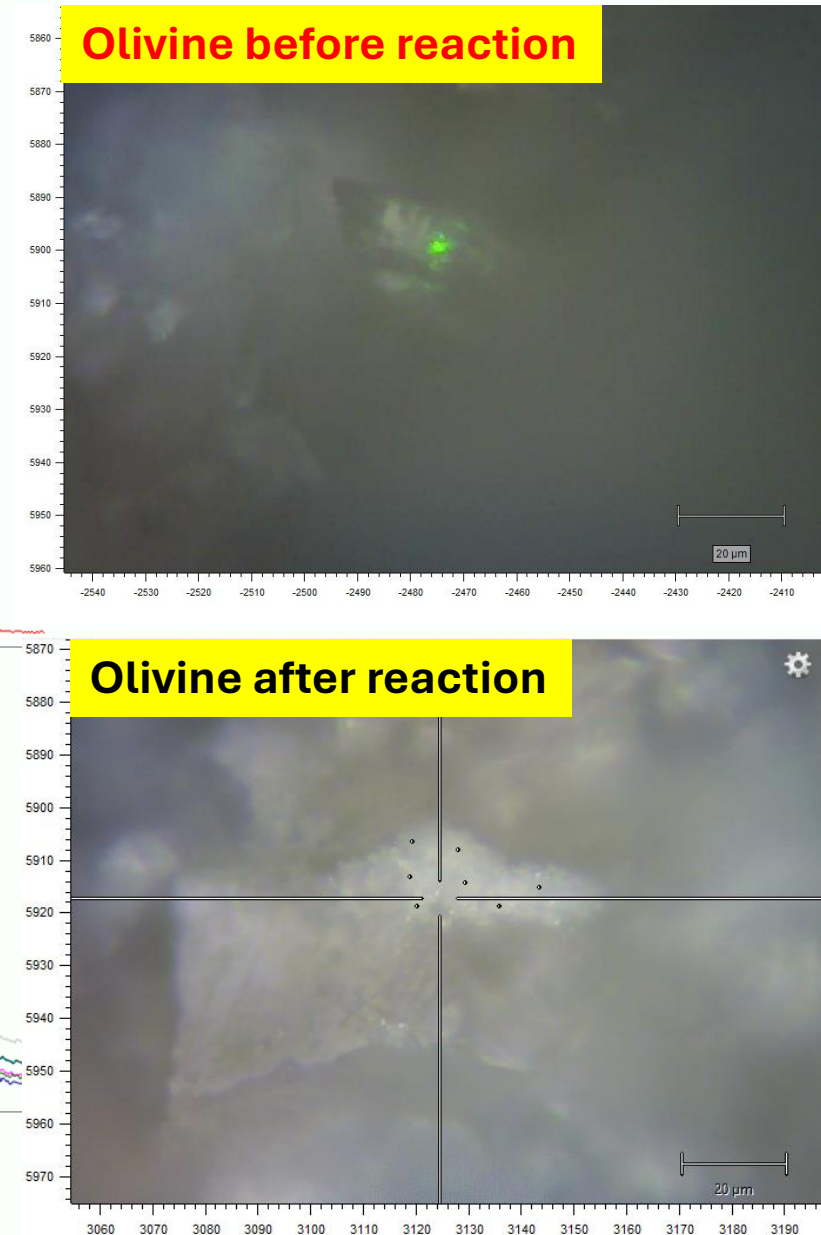
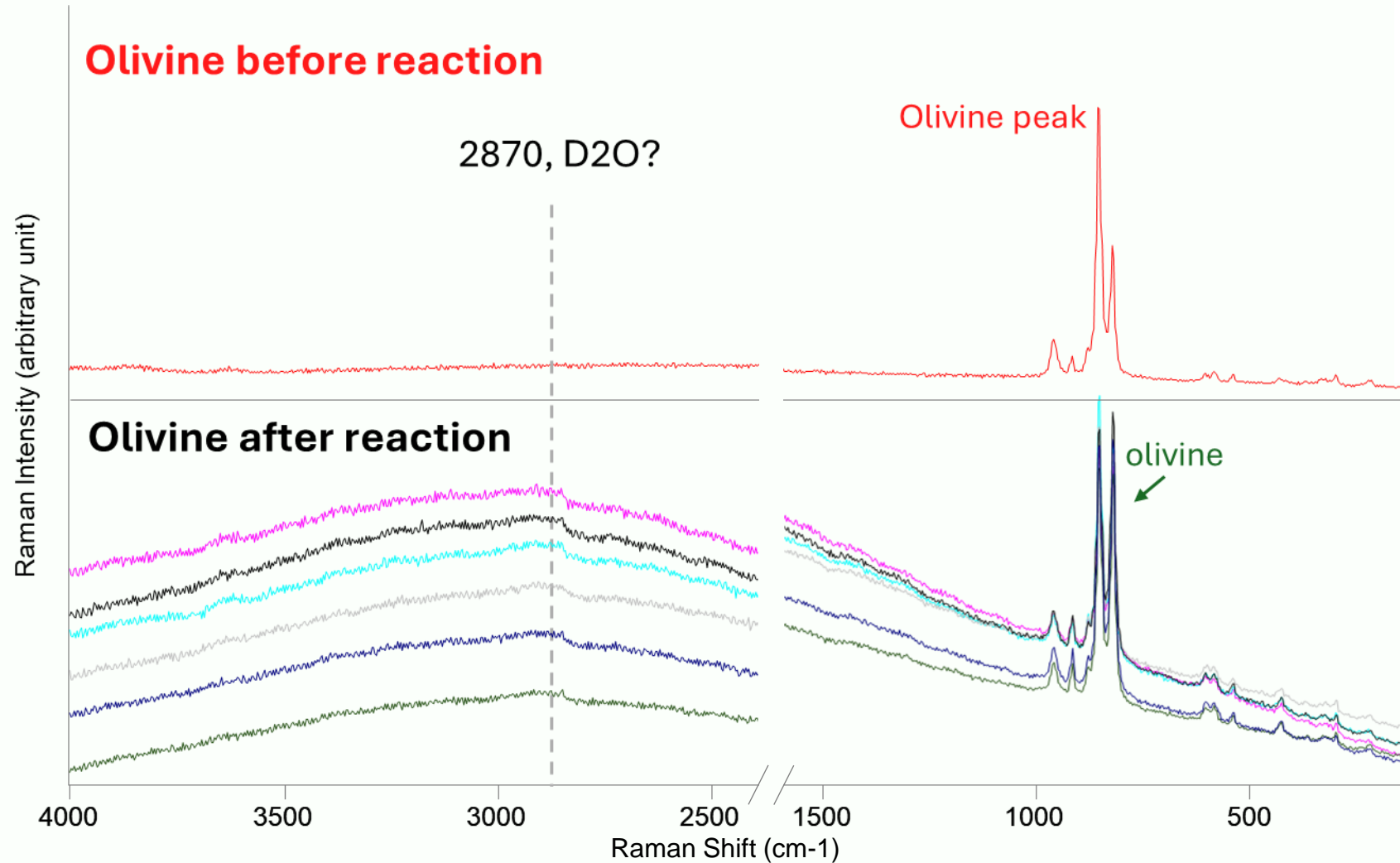
Sample (olivine)

Glove box (N₂ gas)

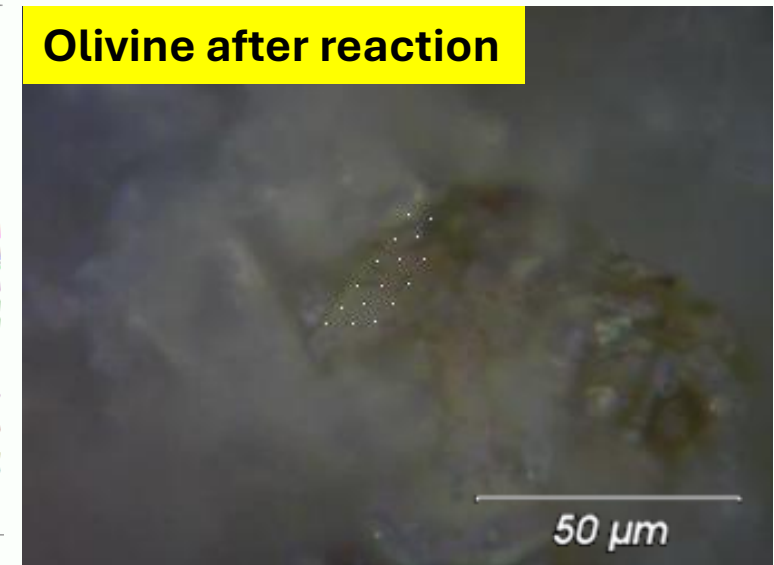
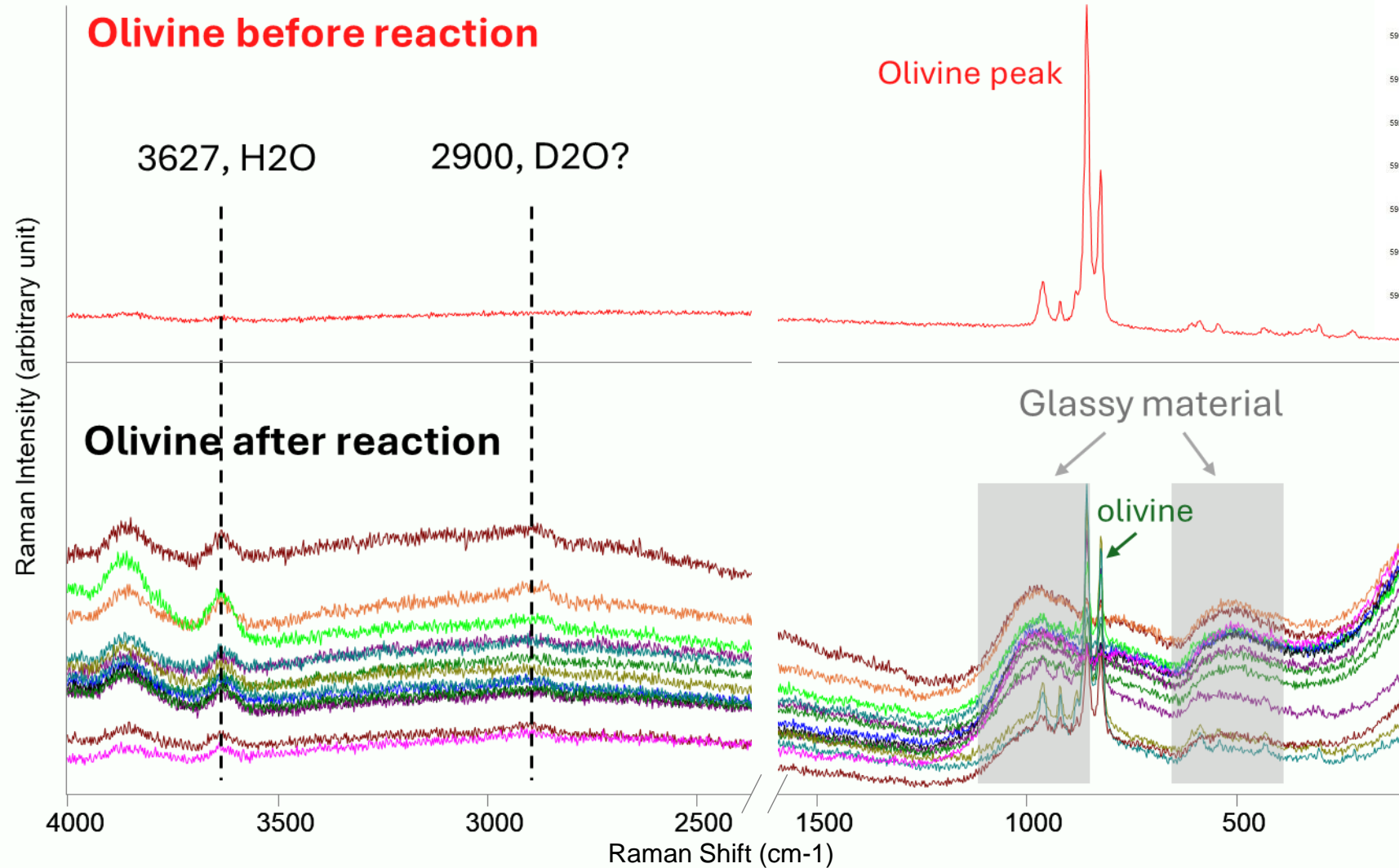
Vacuum chamber (Ar gas)



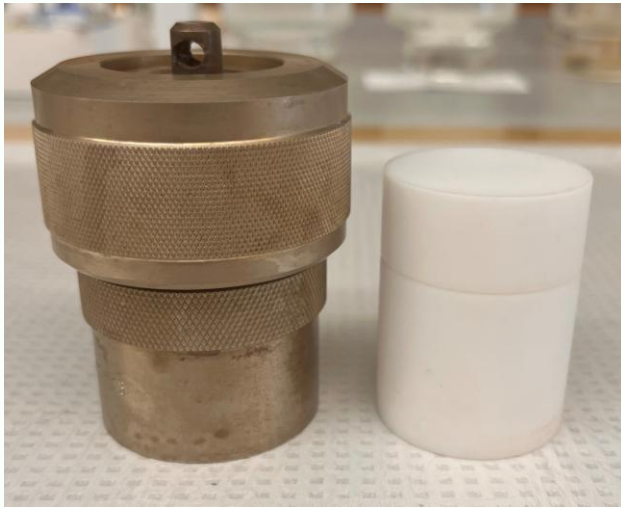
Exp 1: D₂O feature after reaction



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Experiment 2: Pressurized reaction vessel



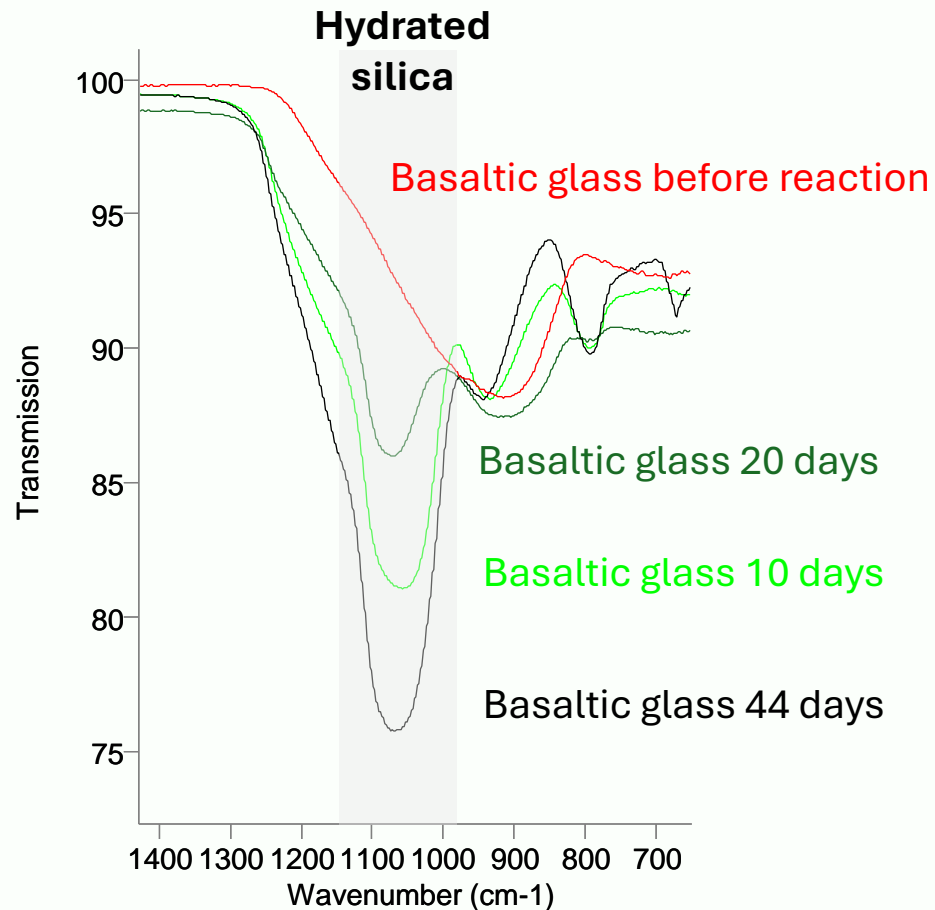
- Teflon-lined
- Solid: olivine or basaltic glass 50 mg
- Solution: 0.5 ml D₂O + 0.5 ml dilute H₂SO₄ (16%)
- Temperature: 90 °C
- Duration: 10 days, 20 days, 44 days
- Characterization: FT-IR and Raman

Bulk composition
change

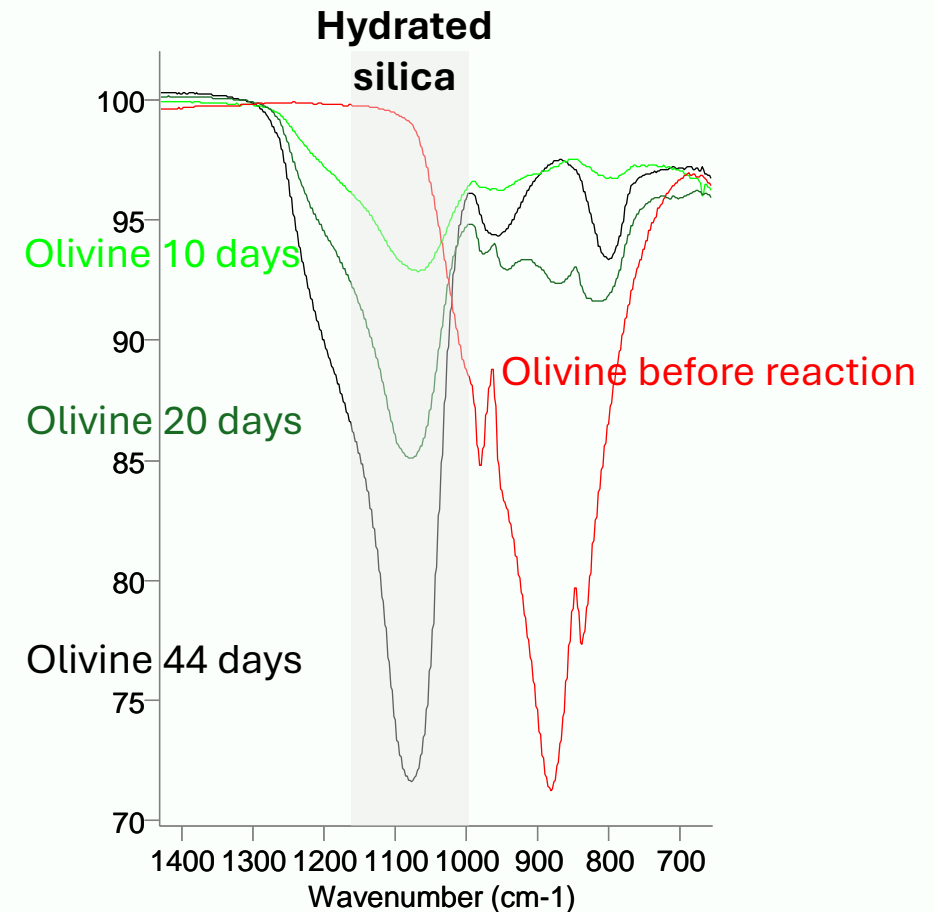
Micro-analysis on
single points

Exp 2: FT-IR spectra – the main reaction product is hydrated silica

Basaltic glass IR spectra

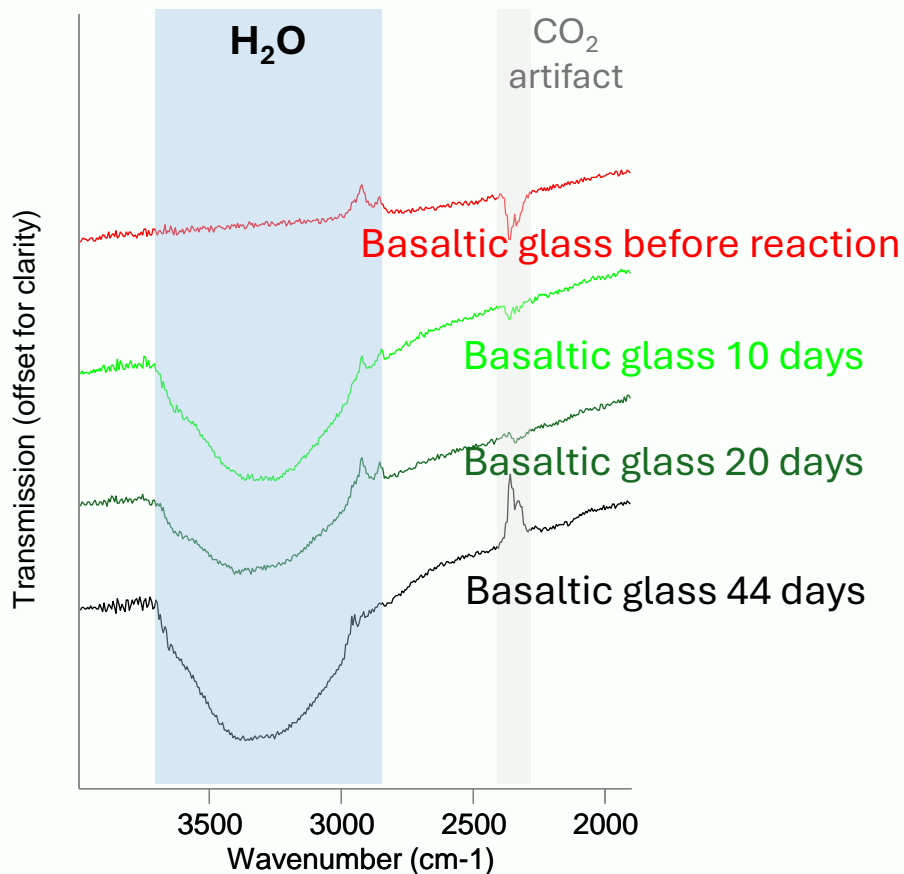


Olivine IR spectra

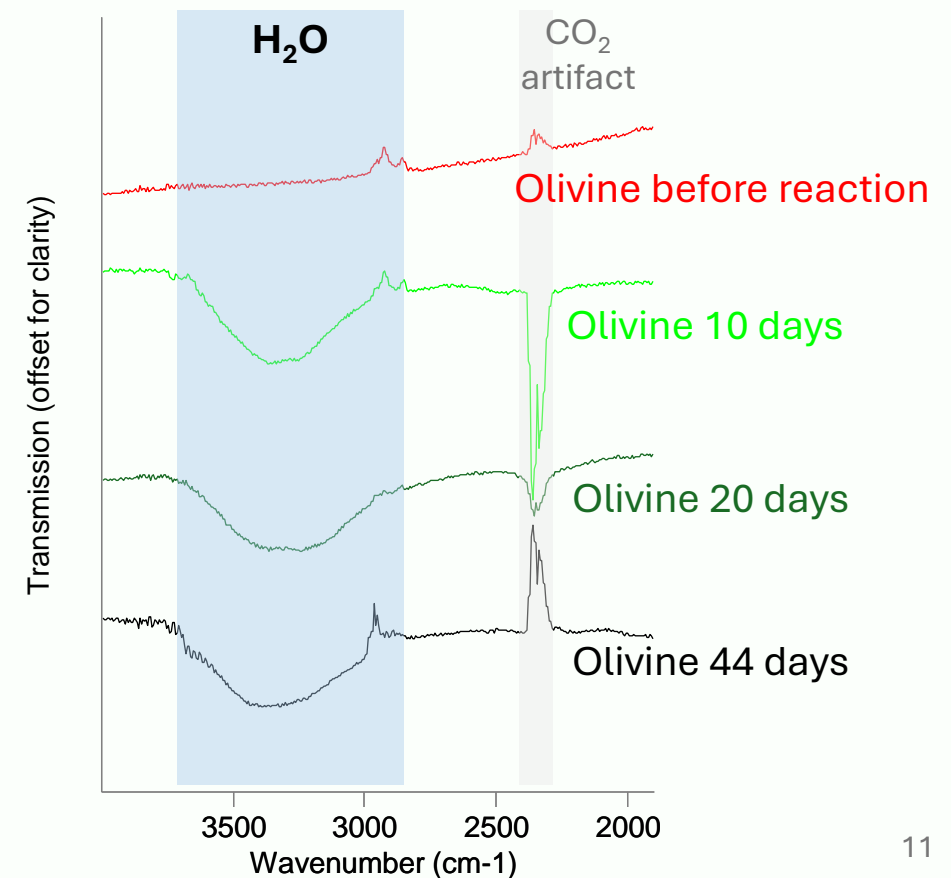


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Olivine IR spectra



Experiment 2: Pressurized reaction vessel

Summary of Raman features found after different reactions

Duration	Olivine	Basaltic glass
10 days	Sulfate feature	No apparent change
20 days	Hematite feature	No apparent change
44 days	Completely dissolved	Sulfate+D ₂ O+H ₂ O

- Olivine: different reaction products depending on reaction time.
- Basaltic glass: dissolves slower than olivine

Experiment 2: Pressurized reaction vessel

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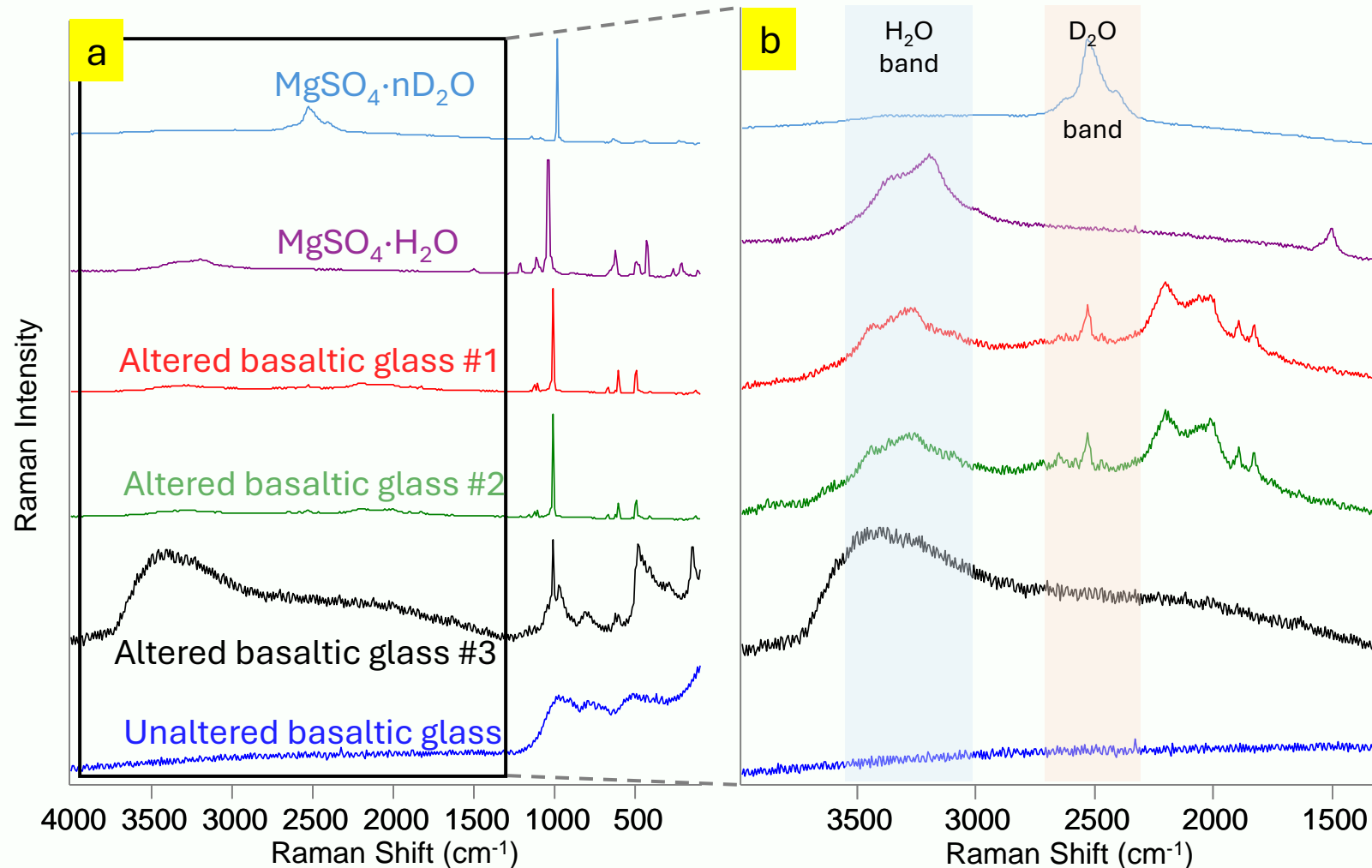
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Exp 2: H₂O, D₂O, sulfate Raman features after 44 days of reaction with basaltic glass



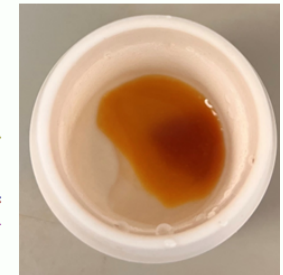
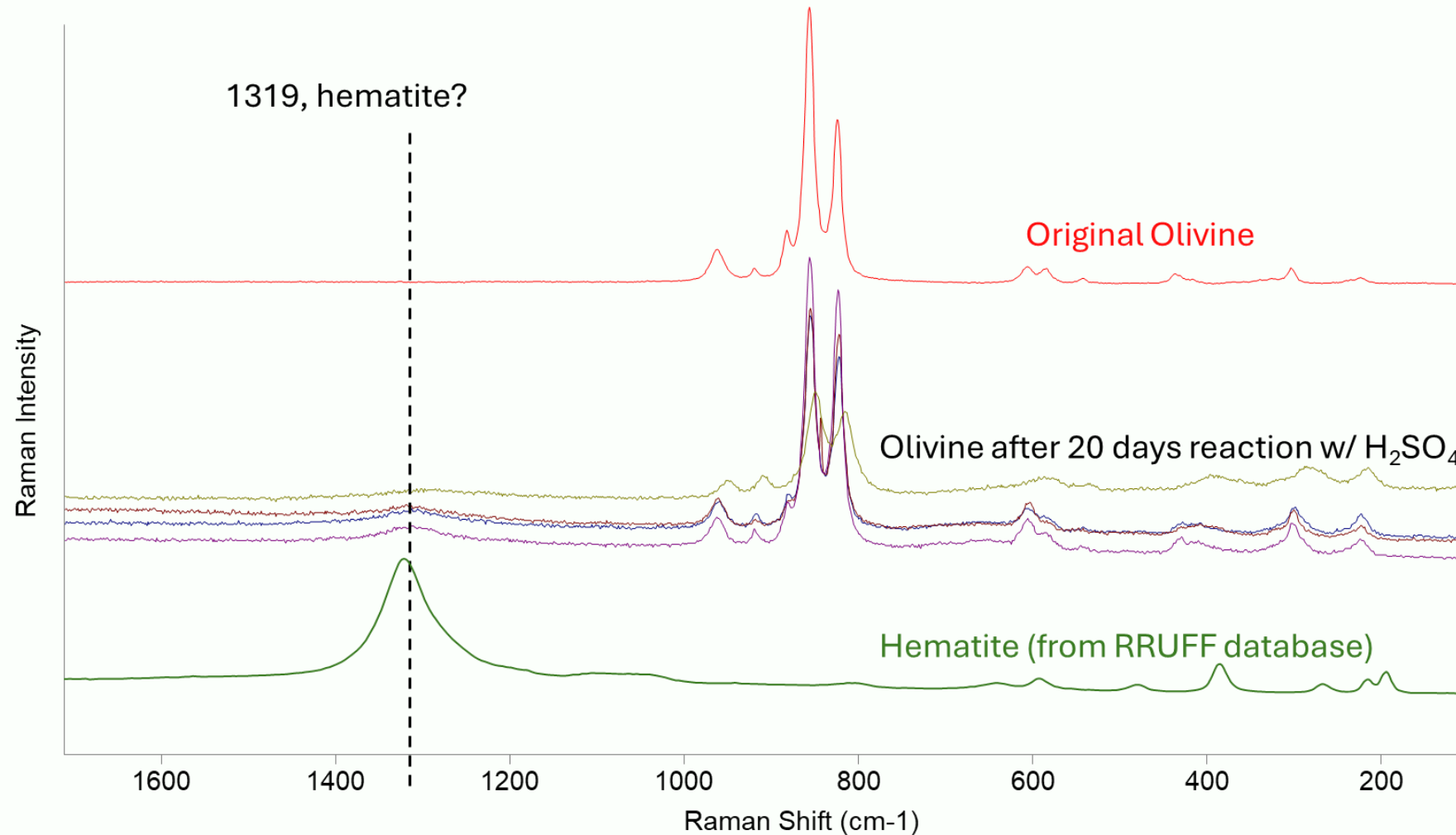
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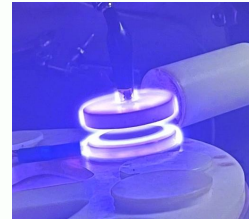
Exp 2: Hematite Raman feature after 20 days of reaction in olivine



Conclusions

Exp 1 (EEI + D₂O + heating)

- D₂O feature identified by Raman spectroscopy



EEI effectively activates grain surfaces as we might expect occurs on the lunar surface by space weathering

Exp 2 (Acid + D₂O + longer term heating)

- Multiple alteration products (silica, ferric oxide, sulfate, D₂O enters mineral structure)



- Dissolution of olivine/basaltic glass
- Evaporation of "brine" produces secondary deposits of salts

Thank you!

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